

Scientific Computing with JAX

A Case Study Evaluating Gravitational Lensing Likelihood
[HTML presentation](#), [PDF archive](#)

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June 4th, 2025

Abstract

Presented at [Durham HPC Days 2025](#).

JAX is a Python library that combines just-in-time (JIT) compilation with automatic differentiation, powered by XLA (Accelerated Linear Algebra compiler), to target multiple hardware architectures including CPU, GPU (NVIDIA, AMD, Intel, Apple), and Google TPU. While primarily designed for machine learning research, JAX presents compelling advantages for scientific computing in the HPC landscape.

As HPC increasingly depends on heterogeneous accelerators, JAX’s “write once, deploy anywhere” approach enables scientific libraries to efficiently utilize diverse computing resources. Python programmers familiar with array programming and other JIT solutions like Numba can transition to JAX with minimal code changes, allowing straightforward adoption. Additionally, automatic differentiation provides gradient calculations for optimization problems with minimal developer effort — automatically generating derivatives that would be prohibitively complex to implement manually, while enabling optimization algorithms to converge more rapidly with gradient information.

This presentation examines the adaptation of a likelihood function calculation from Numba to JAX in the context of gravitational lensing analysis of James Webb Space Telescope observations. The case study specifically addresses the convergence of AI/ML techniques with traditional scientific computing, highlighting performance gains, implementation challenges, and practical considerations when migrating existing scientific code to JAX’s programming model. We’ll discuss how JAX enables scalable processing across different HPC hardware resources while maintaining scientific accuracy and computational efficiency.

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Motivations

Physics: revealing the nature of dark matter with the James Webb Space Telescope (JWST)

- Probing the low-mass region ($\lesssim 10^{8.5}M_{\odot}$) of Dark Matter substructures
 - Λ CDM vs. alternative Dark Matter models
 - invisible \Rightarrow probing via strong gravitational lensing
 - demonstrated in James W. Nightingale et al.¹ using *Hubble Space Telescope* (HST) data
- Next: take advantage of 4+ wavebands, high quality observations from *James Webb Space Telescope* (JWST)

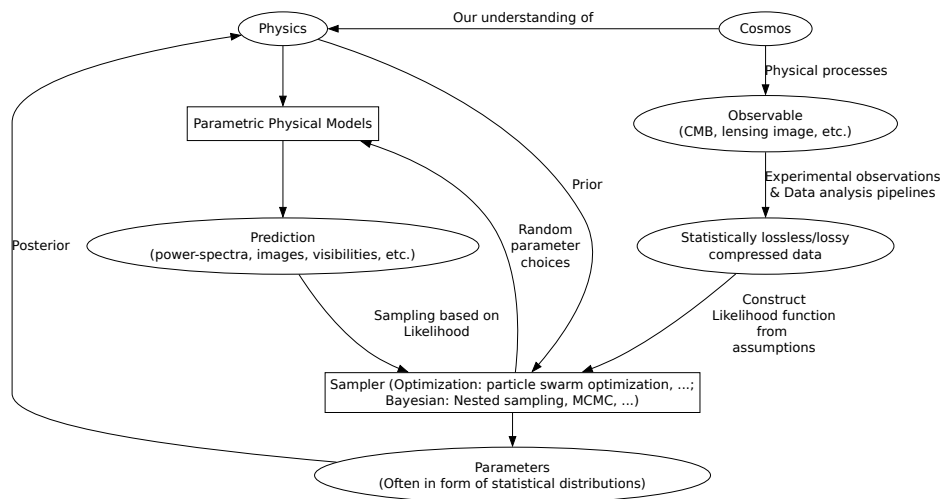
How? PyAutoLens

1. Mass modeling: parametric, non-linear, multi-phase models
2. Source reconstruction: linear, reconstructing unlensed light distribution via different kinds of meshes: rectangular, Delaunay, Voronoi

Log-likelihood function takes the output of (1) and computes its likelihood (where (2) is part of the calculation).

Key goal is to automate the whole process and apply it to large datasets.

The power of likelihood in theory



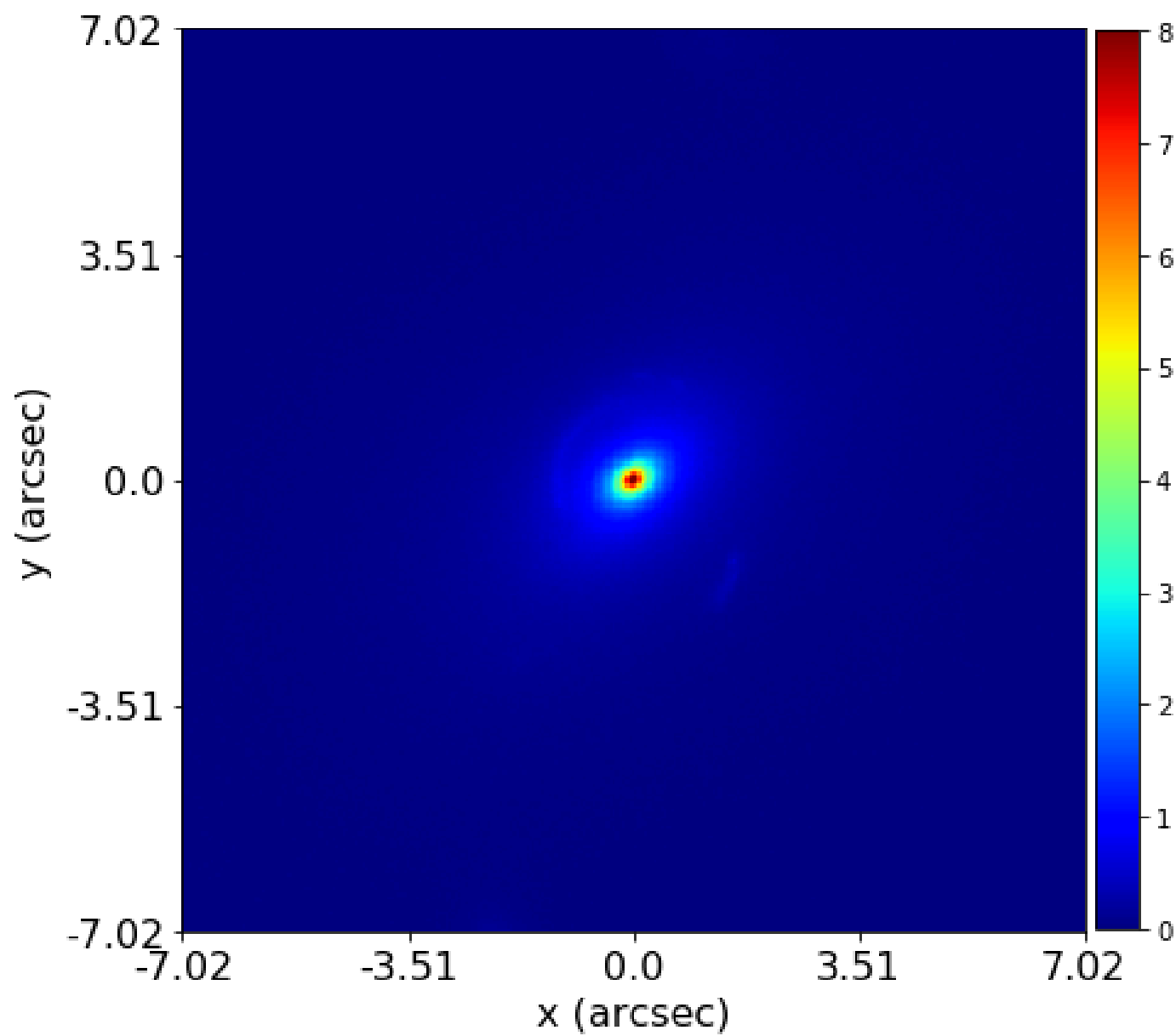
$$\text{Figure 1: } P(\Theta|\mathbf{D}, M) = \frac{P(\mathbf{D}|\Theta, M)P(\Theta|M)}{P(\mathbf{D}|M)} \equiv \frac{\mathcal{L}(\Theta)\pi(\Theta)}{\mathcal{Z}}$$

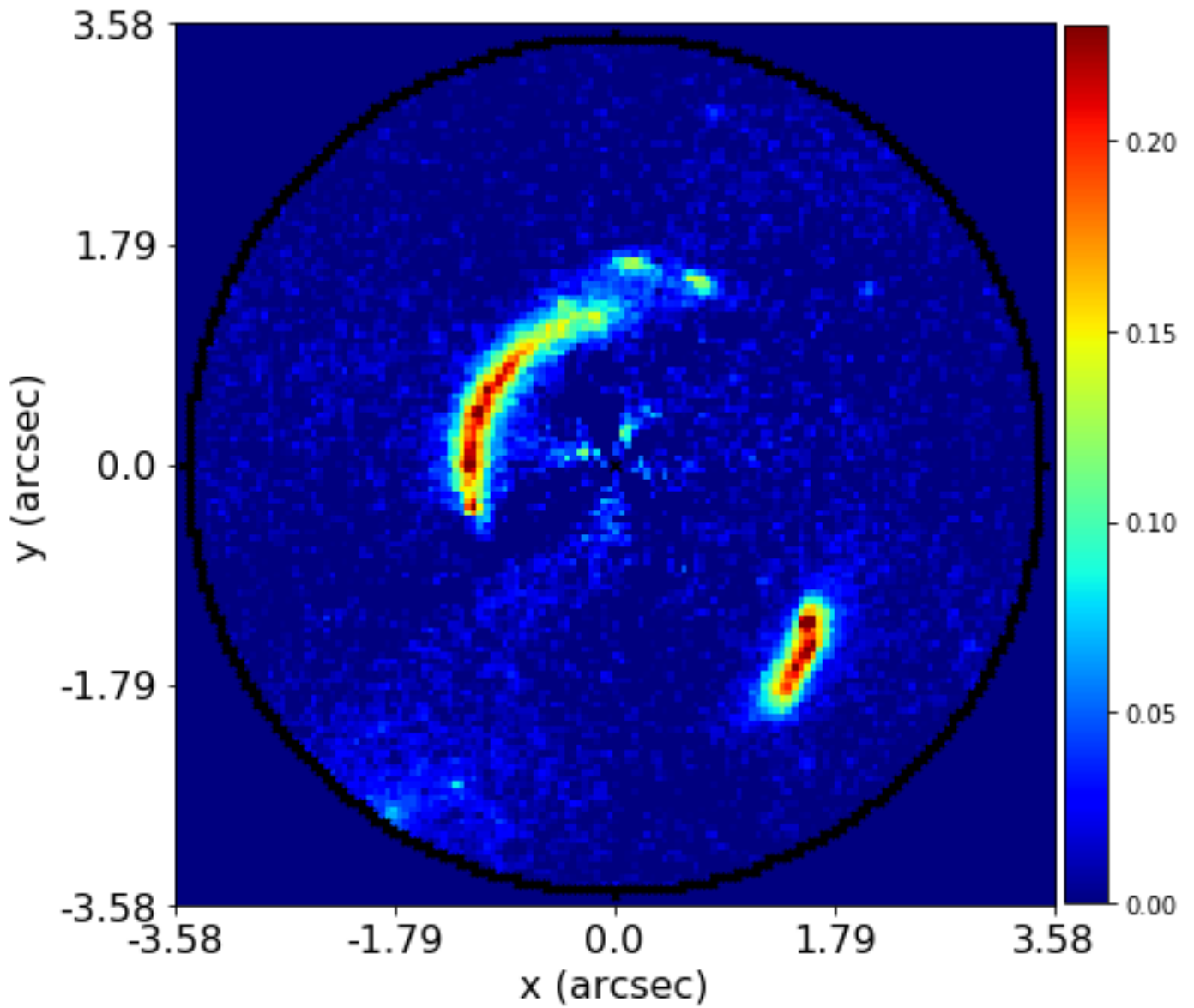
The power of likelihood in action

- 25 free parameters
 - Lens Light (11): Sersic + Exponential
 - Lens Mass (7): SIE + Shear
 - Source Light (7): Sersic

PyAutoLens (via PyAutoFit) supports Nested sampling (**Dynesty**), MCMC (emcee), particle swarm optimization (PySwarms)

¹“Scanning for Dark Matter Subhaloes in *Hubble Space Telescope* Imaging of 54 Strong Lenses,” *Monthly Notices of the Royal Astronomical Society* 527, no. 4 (2023): 10480–506, <https://doi.org/10.1093/mnras/stad3694>.

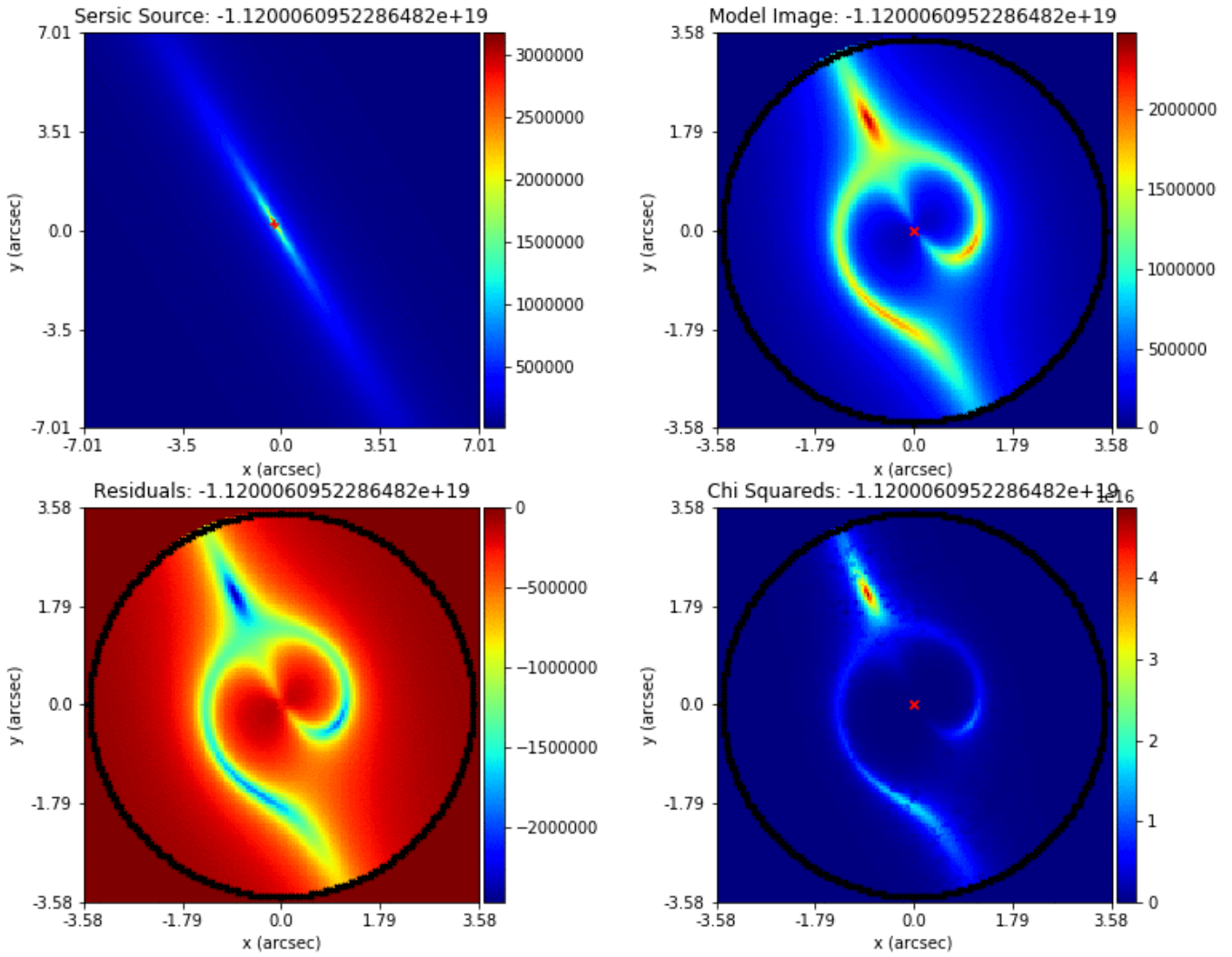




The power of likelihood in action

- 25 free parameters
 - Lens Light (11): Sersic + Exponential
 - Lens Mass (7): SIE + Shear
 - Source Light (7): Sersic

PyAutoLens (via PyAutoFit) supports Nested sampling (**Dynesty**), MCMC (emcee), particle swarm optimization (PySwarms)



Computational challenges

- PyAutoLens originally is implemented in Numba
- For single-band HST data, the image processing analysis of a single lens takes approximately 48 hours over 76 CPUs.
 - image pixels: $M \sim 10,000$
 - source image pixels: $S \sim 2,000$
 - 3 s for 1 iteration, $O(100,000)$ iterations needed.
- For JWST's 4x freq. bands, $M \rightarrow 4M, S \rightarrow 4S$. Runtime: $\sim 3 \text{ s} \rightarrow 60 \text{ s}$. I.e. $\sim \times 20$

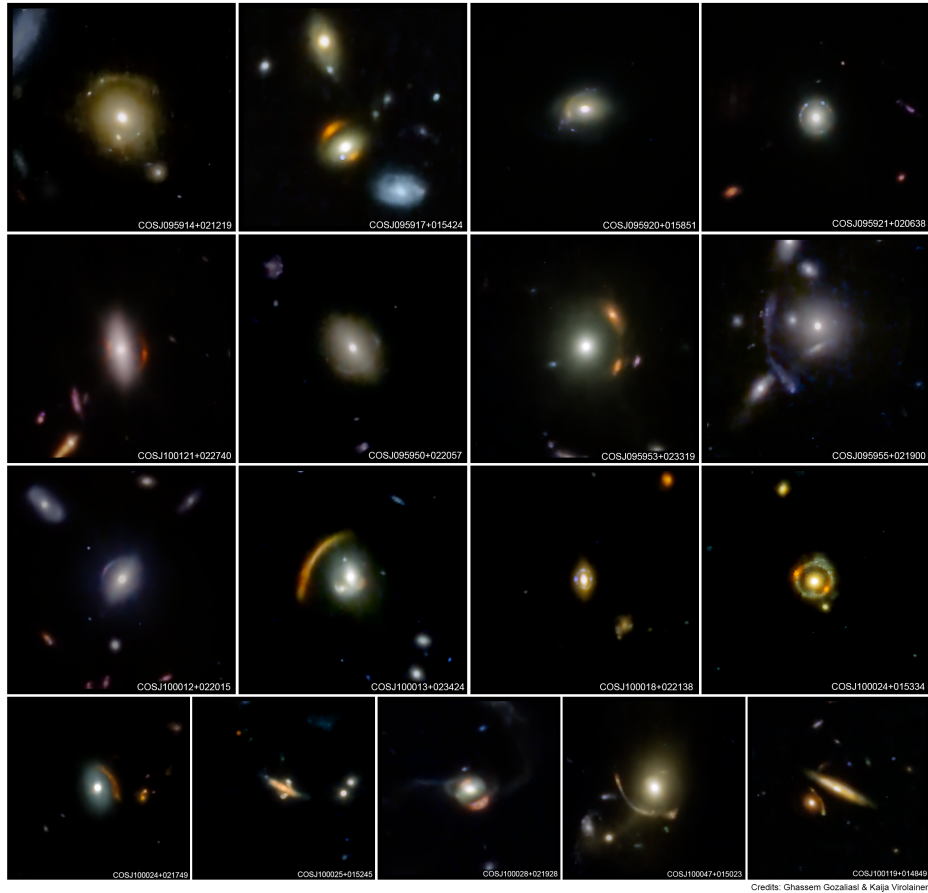


Figure 2: The 17 most spectacular lenses from the COWLS sample, revealed by the JWST imaging through our visual inspection of the COSMOS-Web field. The images are produced combining the four filters (F115W, F150W, F277W, F444W) for an ideal rendering of the lensing evidence.²

Why JAX?

- Herculens,³ GIGA-Lens⁴ demonstrated successful adoption of JAX in modeling strong gravitational lensing.
- Performance gains potentially come from:
 - Functions implemented with JAX become faster
 - Running on accelerators
 - Gradient information reduces no. of iterations in fitting
- This project ports the likelihood function, a subset of functionality provided in PyAutoLens, from Numba to JAX

Lesson learnt (programming experience)

Methodology

- Porting
 1. translate the function to Math
 2. translate the Math to Numba, using vectorized programming as much as possible to anticipate the programming paradigm in JAX
 3. translate the Numba function to JAX, where in simple case would just work
 4. further optimize from there
- Organizing & Testing
 - Functions are then kept in 3 different modules, original for the original functions, numba for those ported in (2), jax for those ported in (3)

²Guillaume Mahler et al., “The COSMOS-Web Lens Survey (COWLS) II: Depth, Resolution, and NIR Coverage from JWST Reveal 17 Spectacular Lenses,” arXiv:2503.08782, preprint, arXiv, March 11, 2025, <https://doi.org/10.48550/arXiv.2503.08782>.

³A. Galan et al., “Using Wavelets to Capture Deviations from Smoothness in Galaxy-Scale Strong Lenses,” *Astronomy & Astrophysics* 668 (December 2022): A155, <https://doi.org/10.1051/0004-6361/202244464>.

⁴A. Gu et al., “GIGA-Lens: Fast Bayesian Inference for Strong Gravitational Lens Modeling,” *The Astrophysical Journal* 935, no. 1 (2022): 49, <https://doi.org/10.3847/1538-4357/ac6de4>.

- Metaprogramming is used in setting up unit-test framework (via `pytest`) to guarantee correctness. `pytest-benchmark` is used to compare performance differences between these implementations. This feeds back into step (4).

What is Numba

- Numba is a jit compiler supporting a subset of Python and NumPy operations, powered by LLVM. While it is possible to target the GPU via CUDA, it requires rewriting the function using different APIs and paradigms, not to mention it is CUDA (i.e. NVidia) only.

What is JAX

- JAX is a jit compiler, tracing compiler by Google, powered by XLA compiler, originated from Google. JAX is designed primarily for machine learning workloads but is suitable for scientific computing as well. It is a tracing compiler removing side-effects of function. I.e. effectively it encourages functional programming paradigm and thinking. It automatically targets multiple hardware architectures including CPU, GPU, TPU, without requiring rewriting.
- I.e. it solves the “two-language problem”, or more accurately, “three-implementation problem”: prototype/API, CPU, GPU.

Numba vs. JAX

- Numba and JAX are both Domain-Specific Languages (DSLs), with different kinds of fallbacks when complete jit compilation of a function is not possible.
 - Better think of it as language + compiler + library.

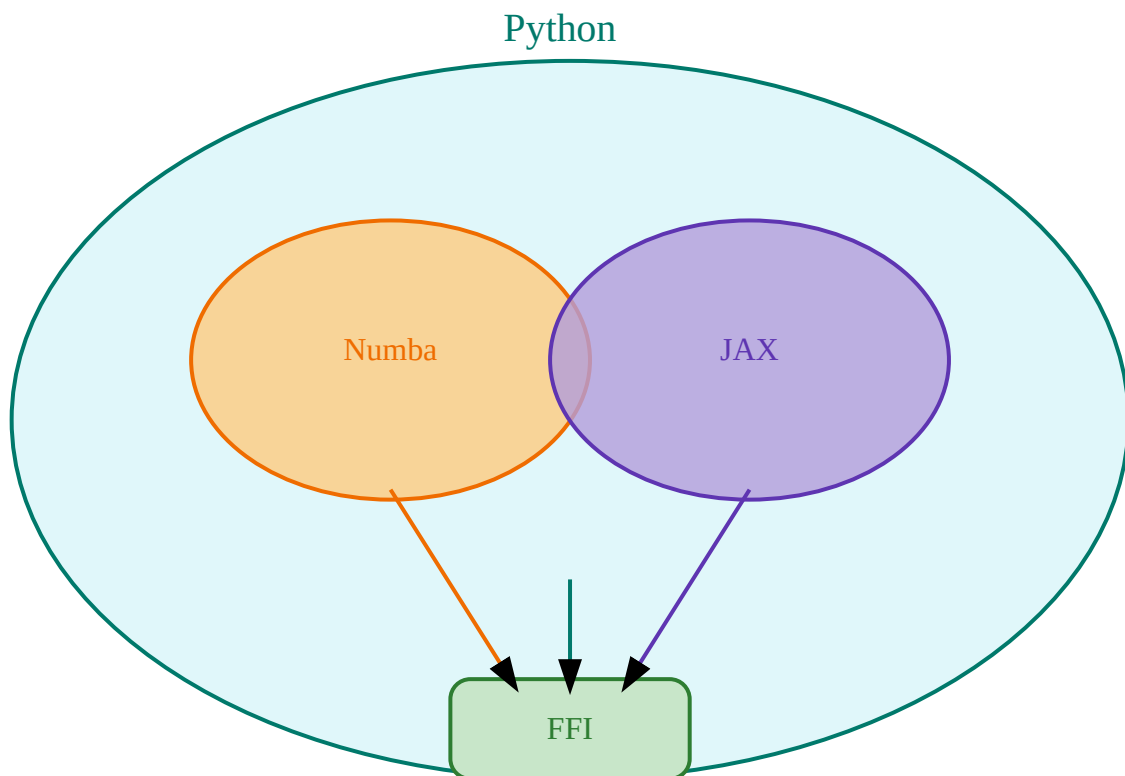


Table 1: Numba vs. JAX

Numba	JAX
C-like mini language	Smaller language ($JAX \subset Numba$): restrictions on control flow, mutation, and dynamic shapes
Implements a subset of Python+NumPy, with a parallelization model similar to a mini-“OpenMP” NumPy implementations are dropped in replacement but only a subset is implemented. Calling NumPy within jitted function is completely hijacked. Documentation is minimal.	Implements a subset of Python+NumPy+SciPy exposed via duck-typing. <code>jax.numpy</code> and <code>jax.scipy</code> have similar API comparing to NumPy and SciPy, but has its own documentation. This facilitates deviations in behaviors .
Functions “recompile” whenever input type changes.	Functions “recompile” whenever input type and shape changes.
No automatic compiling & offloading to accelerator. No autograd/autodiff.	Going through FFI is more costly: memory transfer from and to device, losing autograd/autodiff.

Characteristics of JAX

- tracing compiler & recompile per shape change \Rightarrow `static_argnums`

```
@partial(jax.jit, static_argnums=0)
def this_recompile_everytime(shape):
    return jax.numpy.zeros(shape)
```

- Compiler Driven Design
 - Especially in JAX, partly because of its functional paradigm, framing your problem in JAX idiomatic expressions results in great speed up, sometimes more than you could do otherwise in Numba because of its design (recompile per shape, fusion/fusing compatible operations, etc.), but you’ll hit a wall if you want more low-level optimizations.
 - It can also means performance improvements can come for free through compiler improvements, as long as your code is written in JAX idiomatic way.
- Easy to port to GPU without setting one up.
- JAX vs numba-cuda: The XLA compiler handles device-specific optimization automatically.
- JAX nudges you to write correct code, and performance comes as a bonus.

Benchmark analysis

\tilde{w} —Code: original version

$$\tilde{W}_{ij} = \sum_{k=1}^K \frac{1}{n_k^2} \cos(2\pi[(g_{i1} - g_{j1})u_{k0} + (g_{i0} - g_{j0})u_{k1}])$$

```
@numba.jit(nopython=True, nogil=True, parallel=True)
def w_tilde_curvature_interferometer_from(
    noise_map_real: np.ndarray,
    uv_wavelengths: np.ndarray,
    grid_radians_slim: np.ndarray,
) -> np.ndarray:
    w_tilde = np.zeros((grid_radians_slim.shape[0], grid_radians_slim.shape[0]))

    for i in range(w_tilde.shape[0]):
        for j in range(i, w_tilde.shape[1]):
            y_offset = grid_radians_slim[i, 1] - grid_radians_slim[j, 1]
            x_offset = grid_radians_slim[i, 0] - grid_radians_slim[j, 0]

            for vis_ld_index in range(uv_wavelengths.shape[0]):
                w_tilde[i, j] += noise_map_real[vis_ld_index] ** -2.0 * np.cos(
                    2.0
                    * np.pi
                    * (y_offset * uv_wavelengths[vis_ld_index, 0] + x_offset *
                      uv_wavelengths[vis_ld_index, 1])
```

```

    )

    for i in range(w_tilde.shape[0]):
        for j in range(i, w_tilde.shape[1]):
            w_tilde[j, i] = w_tilde[i, j]

    return w_tilde

```

\tilde{w} —Code: 1st try

$$\tilde{W}_{ij} = \sum_{k=1}^K \frac{1}{n_k^2} \cos(2\pi[(g_{i1} - g_{j1})u_{k0} + (g_{i0} - g_{j0})u_{k1}])$$

```

@jax.jit
def w_tilde_curvature_interferometer_from(
    noise_map_real: np.ndarray[tuple[int], np.float64],
    uv_wavelengths: np.ndarray[tuple[int, int], np.float64],
    grid_radians_slim: np.ndarray[tuple[int, int], np.float64],
) -> np.ndarray[tuple[int, int], np.float64]:
    # (M, M, 1, 2)
    g_ij = grid_radians_slim.reshape(-1, 1, 1, 2) - grid_radians_slim.reshape(1, -1, 1, 2)
    # (1, 1, K, 2)
    u_k = uv_wavelengths.reshape(1, 1, -1, 2)
    return (
        jnp.cos(
            (2.0 * jnp.pi) *
            # (M, M, K)
            (
                g_ij[:, :, :, 0] * u_k[:, :, :, 1] +
                g_ij[:, :, :, 1] * u_k[:, :, :, 0]
            )
        ) /
        # (1, 1, K)
        jnp.square(noise_map_real).reshape(1, 1, -1)
    ).sum(2) # sum over k

```

\tilde{w} —Code: 2nd try

$$\tilde{W}_{ij} = \sum_{k=1}^K \frac{1}{n_k^2} \cos(2\pi[(g_{i1} - g_{j1})u_{k0} + (g_{i0} - g_{j0})u_{k1}])$$

```

@jax.jit
def w_tilde_curvature_interferometer_from(
    noise_map_real: np.ndarray[tuple[int], np.float64],
    uv_wavelengths: np.ndarray[tuple[int, int], np.float64],
    grid_radians_slim: np.ndarray[tuple[int, int], np.float64],
) -> np.ndarray[tuple[int, int], np.float64]:
    # A_mk, m<M, k<K
    # assume M > K to put TWO_PI multiplication there
    A = grid_radians_slim @ (TWO_PI * uv_wavelengths[:, ::-1]).T

    noise_map_real_inv = jnp.reciprocal(noise_map_real)
    C = jnp.cos(A) * noise_map_real_inv
    S = jnp.sin(A) * noise_map_real_inv

    return C @ C.T + S @ S.T

```

\tilde{w} —Code: digression in problem sizes

$$\tilde{W}_{ij} = \sum_{k=1}^K \frac{1}{n_k^2} \cos(2\pi[(g_{i1} - g_{j1})u_{k0} + (g_{i0} - g_{j0})u_{k1}])$$

Number of image pixels $M \sim 70,000 \Rightarrow M^2 \sim 5 \times 10^9$, $0 \leq i, j < M$
 $N \sim \sqrt{M} \sim 300$

Number of visibilities $K \sim 10^7$, $0 \leq k < K$

$(M, M, K, 2)$ of 64-bit array would be ~ 700 PiB!

While (M, M) of 64-bit array would be ~ 40 GiB only.

\tilde{w} —Code: final try—Numba

$$\tilde{W}_{ij} = \sum_{k=1}^K \frac{1}{n_k^2} \cos(2\pi[(g_{i1} - g_{j1})u_{k0} + (g_{i0} - g_{j0})u_{k1}])$$

```
@numba.jit("f8[:, ::1](f8[:, ::1], f8[:, ::1], f8[:, ::1]", nopython=True, nogil=True, parallel=True)
def w_tilde_curvature_interferometer_from(
    noise_map_real: np.ndarray[tuple[int], np.float64],
    uv_wavelengths: np.ndarray[tuple[int, int], np.float64],
    grid_radians_slim: np.ndarray[tuple[int, int], np.float64],
) -> np.ndarray[tuple[int, int], np.float64]:
    M = grid_radians_slim.shape[0]
    K = uv_wavelengths.shape[0]
    g_2pi = TWO_PI * grid_radians_slim
    dg_2pi = g_2pi.reshape(-1, 1, 2) - g_2pi.reshape(1, -1, 2)

    w = np.zeros((M, M))
    for k in numba.prange(K):
        w += np.cos(dg_2pi[:, :, 1] * uv_wavelengths[k, 0] + dg_2pi[:, :, 0] * uv_wavelengths[k, 1]) *
        np.reciprocal(
            np.square(noise_map_real[k])
        )
    return w
```

\tilde{w} —Code: final try—JAX

```
@jax.jit
def w_tilde_curvature_interferometer_from(
    noise_map_real: np.ndarray[tuple[int], np.float64],
    uv_wavelengths: np.ndarray[tuple[int, int], np.float64],
    grid_radians_slim: np.ndarray[tuple[int, int], np.float64],
) -> np.ndarray[tuple[int, int], np.float64]:
    M = grid_radians_slim.shape[0]
    g_2pi = TWO_PI * grid_radians_slim
    dg_2pi = g_2pi.reshape(M, 1, 2) - g_2pi.reshape(1, M, 2)
    dg_2pi_y = dg_2pi[:, :, 0]
    dg_2pi_x = dg_2pi[:, :, 1]

    def f_k(
        noise_map_real: float,
        uv_wavelengths: np.ndarray[tuple[int], np.float64],
    ) -> np.ndarray[tuple[int, int], np.float64]:
        return jnp.cos(dg_2pi_x * uv_wavelengths[0] + dg_2pi_y * uv_wavelengths[1]) * jnp.reciprocal(
            jnp.square(noise_map_real)
        )

    def f_scan(
        sum_: np.ndarray[tuple[int, int], np.float64],
        args: tuple[float, np.ndarray[tuple[int], np.float64]],
    ) -> tuple[np.ndarray[tuple[int, int], np.float64], None]:
        noise_map_real, uv_wavelengths = args
        return sum_ + f_k(noise_map_real, uv_wavelengths), None

    res, _ = jax.lax.scan(
        f_scan,
```

```

jnp.zeros((M, M)),
(
    noise_map_real,
    uv_wavelengths,
),
)
return res

```

Match 1: Numba vs JAX with 1 CPU core

Table 2: $N = 64$, $B = 3$, $K = 32768$, $P = 32$, $S = 256$, $w_tilde_curvature_interferometer_from$

Implementation	s	σ
jax_compact_expanded	2.5535 (1.0)	0.0032
jax_compact	2.5543 (1.00)	0.0050
numba_compact	2.8768 (1.13)	0.0012
numba_compact_expanded	2.8967 (1.13)	0.0004
original_preload	11.0392 (4.32)	0.0010
original_preload_expanded	11.0686 (4.33)	0.0005
jax	3,368.6229 (>1000.0)	1.6803
original	3,561.0805 (>1000.0)	0.2255
numba	3,702.7006 (>1000.0)	0.7385

Match 2: Numba with 128 CPU cores and JAX with CUDA on GPU (A100)

Table 3: $N = 32$, $B = 300$, $K = 8192$, $P = 32$, $S = 256$, $w_tilde_curvature_interferometer_from$

Implementation	ms	σ
numba_compact	2.5029 (1.0)	0.0520
numba_compact_expanded	3.7808 (1.51)	0.0415
jax_compact_expanded	58.6799 (23.44)	9.1624
jax_compact	61.5560 (24.59)	6.8555
jax	143.2451 (57.23)	0.0749
original_preload	840.0727 (335.63)	0.1761
original_preload_expanded	842.6588 (336.67)	0.4933
numba	1,794.1949 (716.83)	14.9648
original	69,304.2738 (>1000.0)	13.5543

Bonus round 1: Numba vs JAX with 1 CPU core (F)

$$F = T^T \tilde{w} T$$

Table 4: $N = 64$, $B = 3$, $K = 32768$, $P = 32$, $S = 256$, $curvature_matrix$

Implementation	ms	σ
numba_sparse	8.0733 (1.0)	0.0501
jax	19.7302 (2.44)	1.6986
jax_sparse	25.0091 (3.10)	0.1484
jax_BC00	48.5340 (6.01)	0.1571
numba_compact_sparse	49.8400 (6.17)	0.0794
original_preload_direct	99.2061 (12.29)	0.3163
numba	125.3019 (15.52)	0.1143
original	132.4863 (16.41)	0.1376
numba_compact_sparse_direct	139.9244 (17.33)	0.1562
jax_compact_sparse_BC00	379.7214 (47.03)	1.4144
jax_compact_sparse	380.8865 (47.18)	2.4322

Bonus round 2: Numba with 128 CPU cores and JAX with CUDA on GPU (A100) (F)

Table 5: $N = 32$, $B = 300$, $K = 8192$, $P = 32$, $S = 256$, curvature_matrix

Implementation	μs	σ
jax	260.5957 (1.0)	29.3714
jax_BC00	3,078.2068 (11.81)	35.9463
jax_compact_sparse_BC00	3,207.3388 (12.31)	107.0798
numba_sparse	5,548.5175 (21.29)	64.3711
jax_compact_sparse	7,190.9015 (27.59)	35.7355
numba	18,187.5003 (69.79)	5,603.6081
original	18,279.9851 (70.15)	6,052.1386
jax_sparse	19,786.7200 (75.93)	42.9344
numba_compact_sparse	32,605.2243 (125.12)	248.8764
numba_compact_sparse_direct	1,362,329.9249 (>1000.0)	1,366.9112
original_preload_direct	25,218,633.7856 (>1000.0)	8,722.4870

Lesson learnt (performance characteristics and expectations)

Takeaway from benchmark analysis

- “Porting Numba to Numba” is faster in many cases
- The only fair fight between Numba and JAX is single CPU core benchmark
- Different algorithms of the same function is faster (even across Numba and JAX) depending on input sizes
 - \Rightarrow keep all implementations \rightarrow profile \rightarrow pick best (per science case per system)

Limitation of JAX on CPU

- JAX for multithreading on the CPU is a rabbit hole. All links below are from GitHub issues. The lack of documentation reflects on the lack of interest in multicore parallelism on the CPU from the primarily machine learning community (JAX from Google, XLA from OpenXLA).

From [JAX running in CPU only mode only uses a single core](#):

This is largely working as intended at the moment. JAX doesn't parallelize operations across CPU cores unless you use explicit parallelism constructs like `pmap`. Some JAX operations (e.g., BLAS or LAPACK) operations have their own internal parallelism.

- In an HPC setting, you may want to use multiple hierarchies of parallelism on the CPU: SIMD + Multi-threading (e.g. OpenMP) + Multi-processing (e.g. MPI). In this case, you'd want to limit the number of CPU cores for multi-threading, often set via `..._NUM_THREADS`. This is very obscure in how to achieve such in JAX:
 - `XLA_FLAGS='--xla_cpu_multi_thread_eigen=false intra_op_parallelism_threads=1'` **was the recommendation. It is now recommended to use `NPROC=1`** to disable multithreading used in Eigen instead. Notice the lack of `JAX_NUM_THREADS` \Rightarrow `NPROC` might have side-effects.

```
export MKL_NUM_THREADS=${NUM_THREADS}
export MKL_DOMAIN_NUM_THREADS="MKL_BLAS=${NUM_THREADS}"
export MKL_DYNAMIC=FALSE

export OMP_NUM_THREADS=${NUM_THREADS}
export OMP_PLACES=threads
export OMP_PROC_BIND=spread
export OMP_DYNAMIC=FALSE

export NUMEXPR_NUM_THREADS=${NUM_THREADS}

export OPENBLAS_NUM_THREADS=${NUM_THREADS}

export NUMBA_NUM_THREADS=${NUM_THREADS}

export NPROC=${NUM_THREADS}
export JAX_NUM_CPU_DEVICES=1
export TF_NUM_INTEROP_THREADS=1
export TF_NUM_INTRAOP_THREADS=${NUM_THREADS}
```

- You may set `JAX_NUM_CPU_DEVICES=${NUM_THREADS}` instead, together with sharding to shard your array to different CPU cores. I.e. OpenMP-like parallelism cannot be achieved. Also, `jax.device_put` requires your array length is divisible by `JAX_NUM_CPU_DEVICES`.
 - Preliminary tests are not promising.

Lessons learnt from Numba vs. JAX

- Does it solve the “3-implementation problem”?
 - prototype: original
 - multithreading on the CPU: numba
 - running on accelerator: jax
- Even if we compare JAX and Numba on equal footing (a single CPU core): sometimes it is faster with JAX (e.g. compiler optimization w.r.t. shape, fusion, etc.) but sometimes it is faster with Numba as JAX language is more restrictive and hence Numba allows expression of more efficient algorithms (recall $JAX \subset Numba$)
- Therefore I think it is advantageous to keep both Numba and JAX implementation. See more from [Should Numba be dropped completely? | AutoJAX Doc](#)

Miscellaneous notes

- Decouple algorithmic development (“library code”) and Python API, e.g. allow end users to choose backends between Numba and JAX. This also facilitate future porting to, say, Julia.
- Open problems: Delaunay/Voronoi mesh is going to be difficult to implement as JAX’s programming model dislikes dynamic shapes.
- Don’t wrestle with the language, especially for DSLs

Conclusions

- Bayesian inference in cosmology is fun!
- JAX is a fun(tional) language to work with
 - ✓ very easy to deploy to accelerators
 - ✗ more restrictive in language features
 - ✗ poor on CPU multithreading

Team, Links, & References

- PyAutoLens Team
 - James W. Nightingale
 - Richard G. Hayes
 - Aristeidis Amvrosiadis
 - Coleman Krawczyk
 - Gokmen Kilic
 - ...
- Link of this project: <https://blog.kolen.dev/python-autojax/>. This is a standalone framework to compare different implementations of functions. Functions are being upstreamed to [PyAutoLens](#).
- References

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Gu, A., X. Huang, W. Sheu, et al. “GIGA-Lens: Fast Bayesian Inference for Strong Gravitational Lens Modeling.” *The Astrophysical Journal* 935, no. 1 (2022): 49. <https://doi.org/10.3847/1538-4357/ac6de4>.

Mahler, Guillaume, James W. Nightingale, Natalie B. Hogg, et al. “The COSMOS-Web Lens Survey (COWLS) II: Depth, Resolution, and NIR Coverage from JWST Reveal 17 Spectacular Lenses.” arXiv:2503.08782. Preprint, arXiv, March 11, 2025. <https://doi.org/10.48550/arXiv.2503.08782>.

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